1	Global Distribution and Inter-annual Variations of Mesospheric and
2	Lower Thermospheric Neutral Wind Diurnal Tide
3	Part 1: Migrating Tide
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24	Abstract
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26	Using the TIMED Doppler interferometer (TIDI) mesospheric and lower thermospheric
27	neutral wind multiyear data set (2002 – 2005) and NCAR TIME-GCM 1.2 annual run
28	results at the TIDI sampling points, we study the migrating diurnal tide global

29 distribution, interannual, and seasonal variations in connection with the mean zonal 30 wind interannual and seasonal variations. Strong quasi-biennial oscillation (OBO) 31 effect on the diurnal tide was observed in the TIDI data and reproduced to a less degree 32 in the TIME-GCM run. The migrating diurnal tide amplitude is larger during the 33 eastward phase of the stratosphere QBO and weaker during the westward phase. 34 Westward mesosphere equatorial mean zonal winds appeared during the eastward phase 35 of the stratosphere QBO (2002 and 2004). The strongest QBO effect on both the 36 migrating diurnal tide and mean zonal winds were observed during the northern spring 37 equinox. We believe that is because both the stratosphere winds and diurnal tide 38 amplitude reach their maximum magnitudes during northern spring equinox. The 39 stratospheric QBO winds apply the maximum filtering effect on the gravity waves, which 40 in turn strongly modulate the diurnal and mean zonal winds at higher altitudes. The TIDI 41 data also exhibit large inter-hemispheric asymmetry. The westward mean zonal winds in 42 the mesosphere appeared to be associated with the enhanced diurnal tide. The TIME-43 GCM 1.2 diurnal tide amplitudes are in general smaller than that observed by TIDI. 44 Limited vertical spatial resolution for the TIME-CGM 1.2 is suggested as the cause. 45 Future improvements are expected with higher spatial resolution in the model. 46

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# 48 **1. Introduction**

49 Diurnal tide is the most prominent dynamics feature of the mesosphere and lower 50 thermosphere (MLT) region. The MLT diurnal tide has both significant migrating and 51 nonmigrating components. The migrating tide is sun synchronous caused by solar UV 52 radiation absorption by stratosphere ozone and troposphere water vapor. Studies of the 53 MLT migrating tides have a long history [e.g., Chapman and Lindzen, 1970; Forbes, 54 1982a; 1982b; Vial, 1989]. Due to strong forcing, the migrating diurnal tide is much 55 stronger than the nonmigrating tide in most cases. 56 Both ground based and satellite observations have revealed many important features of 57 the diurnal tide. Ground based observations have established seasonal and interannual

variations of the MLT diurnal tide at various latitudes [e.g., *Manson et al.*, 1988; 1989;

59	1999: Vincent et a	ıl., 1989	She et al.	. 2004: Yuan et al.	. 2006 1. 1	By combining
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60 observations from various locations, *Pancheva et al.*, [2002] were able to assemble a

- 61 global picture of the diurnal tide. But some large gaps in coverage remain.
- 62

63 Satellite observations such HRDI and WINDII from UARS satellite and TIDI from

64 TIMED satellite have been able to provide a global view of the diurnal tide [*Burrage et* 

65 *al.*, 1995; *Wu et al.*, 1995; *Hays et al.*, 1994; *Morton et al.*, 1993; *McLandress et al.*, 1994;

66 1996; Lieberman et al., 1994; Khattatov et al., 1997a; 1997b; Yudin et al., 1998; Wu et

67 *al.*, 2006; *Killeen et al.*, 2006]. More significantly, satellite observations also were able

to separate the nonmigrating tide from the migrating tide [e.g., *Forbes et al.*, 2003;

- 69 *Oberheide et al.*, 2005; 2006].
- 70

71 Migrating diurnal tide has a strong seasonal variation. The tide reaches large amplitudes

72 during spring and fall equinoxes and smaller values during summer and winter solstices.

73 Both ground and satellite observations show this consistent behavior. Various tide

74 models, such as the Global Scale Wave Model (GSWM) also show the same seasonal

changes in tidal amplitude [*Hagan et al.*, 1999a; 2002; 2003].

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77 Moreover, the diurnal tide also exhibits a strong interannual variation closely related to 78 the stratosphere quasi-biennial oscillation (QBO) [Burrage et al., 1995; Hagan et al., 79 1999b, Lieberman, 1997; Vincent et al., 1998]. Recently using TIMED SABER 80 temperature data, Huang et al. [2006] and Xu et al. [2007] have examined the QBO 81 effects on diurnal tide in temperature. It has long been suggested that the diurnal tide 82 QBO effect is related to the mean zonal wind QBO like variations because of the direct 83 link via the advection terms in the tidal equation [Mayr and Mengel, 2005]. 84 Observations have shown that the migrating diurnal tide increases (decrease) when the 85 equatorial mean zonal winds are westward (eastward) [Hagan et al., 1999b, Mavr and 86 Mengel, 2005]. Nevertheless, it has been hard to pin down the causal relationship 87 between the tidal amplitude variations and mean zonal winds. Hagan et al. [1999] used 88 the observed mean zonal winds to simulate the tidal variations and obtained opposite 89 effect on the tidal amplitude using GSWM, whereas, *McLandress* [2002] was able to

90 obtain consistent migrating diurnal tide amplitude with observed mean zonal winds.

91 McLandress [2002] attributed the tidal amplitude increase to the increase of the

92 meridional gradient in the mean zonal winds. *Mayr and Mengel* [2005] suggested that

93 the vertical gradient in the mean zonal wind could also be important in modulating the

94 diurnal tide.

95

96 It has been shown that gravity wave filtering variation due to the stratosphere QBO of 97 zonal winds can generate QBO with opposite phase in the mesospheric zonal winds 98 [Mayr et al., 1997]. Lieberman [1997] also suggested that the momentum flux of 99 migrating diurnal tide (1, 1) mode contribute to the maintenance of time-mean westward 100 winds in 90 - 105 km range. Perhaps, the momentum flux from (1,1) mode may also be 101 partly responsible for the QBO related variations into the westward winds. Obviously, 102 QBO alters the gravity wave filtering pattern and leads to QBO related variations in the 103 migrating diurnal tide and mean zonal winds at high altitudes. Gravity waves may 104 modulate the tide amplitude directly or indirectly through the mean zonal winds. It may 105 also be possible that the gravity wave may interact with the mean zonal wind indirectly 106 through the tide. Thus, it is unclear that the interaction between the tide amplitude and 107 mean zonal winds is one-way or two-way. It will require more observations and model 108 simulation to resolve these issues.

109

110 In this paper, we use recent mesosphere neutral wind data from the TIMED instrument

111 TIDI in combination with the NCAR TIME-GCM 1.2 model annual-run results to

112 examine the QBO effect on the migrating diurnal tide and mean zonal winds. In a

113 companion paper [Wu et al., 2007, hereafter Part II], we will examine the QBO effect on

114 the nonmigrating diurnal tides. We hope to shed some light on this very complex

relationship between the diurnal amplitude and mean zonal winds under the influence of

116 the QBO.

117

118 Since its launch on December 7, 2001, the TIMED satellite has been in orbit for over 5

119 years. The TIDI instrument is a new generation of spaceborne interferometer, and it

120 provides wider latitudinal and nighttime altitudinal coverage compared to the UARS

121 HRDI instrument. Furthermore, TIDI is a dedicated mesosphere and lower 122 thermosphere instrument with more completed and consistent coverage of the MLT 123 region. Because of the 24 hour local time coverage, the mean zonal wind results are less 124 likely to be affected by diurnal tide aliasing. Of course the 24-hour local time coverage 125 is achieved by combining data sampled over 60 days. We also sampled the TIME-126 GCM 1.2 [Roble and Ridley, 1994] annual run at the TIDI data points and analyzed the 127 TIME-GCM 1.2 results in the same manner. In this way, we eliminate any differences 128 between the observations and model due to sampling differences. The TIME-GCM 1.2 129 model is driven by NCEP data (winds and temperature) at 10 hPa (roughly 30 km). 130 Even though the TIME-GCM does not have the explicit QBO, the NCEP data will 131 contain its effect, particularly in the stratosphere neutral winds. These QBO signatures, 132 in turn, may manifest into the mesosphere and lower thermosphere region. Thus give us 133 an opportunity to study the interaction between the diurnal tide and mean zonal winds. 134 This is probably the first attempt to use the TIME-GCM to study inter-annual variations. 135 136 The paper is organized as follows. In Section 1, we give a brief description of the 137 stratosphere environment for the time period of interest and information about the 138 instrument. Section 2 describes data analysis method. Section 3 shows the analysis

results of the migrating diurnal tide in both the meridional and zonal winds. Section 4 isdevoted to the mean zonal winds. We discuss the results in Section 5 and summarize our

141 findings in the Section 6.

### 142 **2. TIDI Data**

#### 143 **2.1 Background Information**

TIDI data from 2002 day 051 (February 20) to the end of 2005 are used in the study, which roughly cover four year time period. Over that time period, the stratosphere zonal winds have experienced roughly two QBO cycles. Figure 1 shows the stratosphere zonal wind over Singapore from 2002 to 2005. The QBO cycles are apparent. During the years of 2002 and 2004 the stratosphere zonal winds are in the eastward phase, while during 2003 and 2005 in the westward phase. The westward phase winds are usually stronger than that during the eastward phase.

#### 152 **2.2 Brief Instrument Description and Data Sampling Distribution**

153 TIDI instrument is a limb-scan Fabry-Perot interferometer, which monitors the Doppler shift in the airglow emissions due to the neutral winds. The details of the instrument can 154 155 be found in papers by Killeen et al. [1999; 2006] and Skinner at al. [2003]. In short, the 156 instrument samples neutral winds on two sides of the satellite track. While one sampling 157 track on one side reaches the North Pole, the other flies over the South Pole. Because 158 the TIMED satellite precesses slowly in local time (3 deg/day) or 12 hours in 60 days, the 159 local time of the sampling points do not change much during a day. On the other hand, 160 the satellite has ascending and descending nodes, which are separated close to 12 hours. 161 Therefore, the satellite can cover 24 hour in local time for every 60 days with combined 162 ascending and descending node data. The satellite orbits the Earth 15 times per day. 163 Figure 2 plots the local times and latitudes of TIDI sample for day 2004080. The red 164 track goes to the North Pole (northern track) whereas the blue track (southern track) 165 passes over the South Pole. The solid dots are for the daytime and hollowed dots for the 166 nighttime. Over a day, the whole sampling pattern will go over the Earth 15 times 167 above different longitudes without drifting much in local time. Over a 60-day period, 168 the whole sampling pattern will shift in local time by 12 hours. One thing should be 169 noted is overlapping of the sampling region. Because both sides make measurements 170 between 60°S and 60°N, there are four local times coverage in this region. Outside this 171 latitude region, the coverage is limited to one track with two local times. This 172 coverage difference will have an impact on precision of the tidal analysis as we will show 173 later on. In the vertical direction, the TIDI instrument scans from 70 to 115 km during 174 the day and 80 to 105 km during the night. We selected the wind data from 85 to 107.5 175 km in our analysis, the data values at these lower and higher boundaries tend to have 176 larger uncertainties due to the lack of airglow emission from these heights. The NCAR 177 processed O2 (0-0) P9 line TIDI data version 0307 are used in this analysis.

### 178 **3. Data Analysis Method**

Because of TIDI's limited local time coverage during a day, it is impossible to calculatethe tidal amplitude and phase without making assumptions about the latitudinal and

181 vertical variations of the tide. Indeed, one can obtain daily variations of the diurnal and 182 semidiurnal tides by assuming that tides have only the most dominant Hough modes 183 [Hays et al., 1994, Lieberman and Hays, 1994]. In another approach, one would assume 184 nothing about the latitudinal and vertical variations, yet imply that the tide does not 185 change much over a period (60 days in case of TIDI). One can use least square method 186 for spectral analysis of space-time series to extract tidal amplitudes and phases at 187 different latitude and altitude [Wu et al., 1995]. This spectral method was also used to 188 extract quasi-two day wave from the TIDI data [Limparsuvan et al., 2005]. Both methods 189 have their limitations and advantages. In this study, we use the second approach. We 190 analyzed the TIDI data based on a 60-day sliding window. Space-time series spectral 191 analysis is performed to both the southern and northern tracks, respectively, to obtain 192 tidal amplitudes and phases of both migrating and nonmigrating tides. The results from 193 the southern and northern tracks are then combined to form a global view of the tidal 194 amplitudes and phases. The final result will have a horizontal resolution of 6 degree in 195 latitude and a vertical resolution of 2.5 km.

196

197 Based on the random statistical error of the wind measurements, we estimate the 198 precision of the tidal amplitude calculation during spring time of each of the four years 199 (2002 - 2005).The results are shown in Figure 3. The precision has improved from 200 2002 to 2005 due to the reduction of ice on TIDI optics through the two roll maneuvers in 201 2003 and gradual sublimation over the years [Skinner et al., 2003]. The vertical 202 variations are mostly due to the  $O_2(0-0)$  band airglow emission profile. The  $O_2(0-0)$ 203 band peaks near 94 km, consequently, the wind error is smallest at that height. Above 204 105 km, only daytime data are available, hence, the error increases. The latitudinal 205 variations are caused by the distribution of sampling point. As we have pointed out that 206 both the southern and northern tracks cover the region between 60°S and 60°N latitudes 207 resulting more data points and smaller statistical errors. In the high latitude regions (> 60208 degree latitude), only one track passes over with fewer sampling points causing higher 209 statistical errors. While it is easier to compute the statistical errors, systematic errors are 210 much difficult to ascertain. Hence, the tidal amplitudes may contain some systematic 211 biases, which will gradually reduce with improvement of data in the future. At this

- 212 point, it appears that we still have some noticeable bias at high latitude region affecting
- 213 the migrating tide analysis. For that reason, we limit our discussion about mid and low
- 214 latitude regions for migrating tide results. Because of high precision, the analysis result
- should provide mostly reliable relative changes for inter-annual comparisons.
- 216

## **4. Diurnal Tide in Mesospheric Winds**

### 218 **4.1 Migrating Diurnal Tide in Merdional Winds**

# 219 **4.1.1 TIDI Results**

220 The migrating tide amplitudes of meridional winds for each season during the four years 221 from 2002 to 2005 are plotted in Figure 4. To highlight the inter-annual variations we 222 plotted the same season of different years side by side. Figure 4a is for the spring 223 equinox. The time period of 60 day starts from day 51 (February 20). The years on 224 the left side (2002 and 2004) are in the eastward phase of the stratosphere QBO. The 225 right side (2003 and 2005) is for the westward phase. The contrast between the left and 226 right is quite apparent. Stronger diurnal tide amplitudes are observed during the 227 eastward phase of the stratosphere QBO. Besides the inter-annual variations, inter-228 hemispheric differences are also quite noticeable. The amplitude peak altitudes in the 229 two hemispheres have offset in height and the southern hemisphere tends to have double 230 peaks. The large tidal amplitudes at northern high latitude and high altitudes appear to 231 be a result of bias in the winds, which tend to affect the migrating tide results. For this 232 reason, we have limited our analysis of migrating tide to 70 degree latitude. To 233 examine the seasonal variation, we use Figure 4b to show the tidal amplitude variations at 234 95 km for each of the four years. The vertical coordinate marks the beginning of the 60 235 day sliding window upon which the space-time series spectral analysis was performed. 236 For comparison with monthly values a upward shift of 30 days is needed. The 2002 237 data starts at day 51 (02/20/02). Peaks at spring and fall equinoxes are visible. The 238 spring equinox diurnal tide amplitudes are larger during 2002 and 2004. Because of the 239 late start in data set, the increase in 2002 is not as obvious as that in 2004. The 240 amplitude in the northern hemisphere show consistent larger values than that in the southern hemisphere. Figure 4c displays the diurnal tide amplitude at 21° N for the four 241

242 years. The spring equinox maximum is apparent for all four years. The amplitude of the

- 243 migrating diurnal tide appears to shift upward during the spring equinox for all four years.
- Again because of the late start of the 2002 data, the slightly increase in the 2002 spring
- 245 equinox is not clearly seen. We should note that the starting point of 2002 is day 51
- 246 (February 20). On that day we still see the contour of 60 m/s. In 2003, the same 60 m/s
- contour line did not appear to extend to day 51. The increase in 2004 is very obvious.
- 248 There is no clear QBO related variation for the fall equinox peaks.
- 249

## 250 4.1.2 TIME-GCM Results

251 Figure 5a is for the spring equinox results from TIME-GCM annual run. There is a very 252 small increase in the tide amplitude during the eastward phase (2002 and 2004). Both 253 hemispheres show multiple peaks at similar heights with varying amplitudes. The 254 amplitudes are in general smaller than that from the TIDI observations. Figure 5b shows 255 the migrating diurnal tide amplitude at 95 km. We can hardly identify the increase during the spring equinox. Figure 5c illustrates the diurnal tide amplitude at  $21^{\circ}$  N. The 256 257 spring equinox amplitude in 2002 is only slightly larger noting the small 40 m/s contour 258 line near 100 km. The increase is more obvious in 2004. The peak amplitude also 259 appears to shift upward during the spring equinox for all four years.

260

# 261 **4.2 Migrating Diurnal Tide in Zonal Winds**

# 262 **4.2.1 TIDI Results**

263 Figure 6a illustrates the spring equinox zonal wind diurnal amplitudes. The inter-

hemisphere distance between the peak latitudes are larger in the zonal winds than in the

- 265 meridional winds which is consistent zonal wind Hough mode configuration. The
- amplitude tends to peak near  $30^{\circ}$ N in the northern hemisphere. In the southern
- 267 hemisphere, the peak amplitude is much widely distributed into the mid latitude region
- 268 (50°N). We also see a tendency to have larger amplitudes during the eastward phase of
- the stratosphere QBO (2002 and 2004). The amplitude at 95 km also show large spring
- equinox peaks during 2002 and 2004 in the southern hemisphere (Figure 6b).
- 271 4.2.2 TIME-GCM Results

272 The spring equinox amplitudes in the southern hemisphere are larger during the eastward 273 phase of QBO (Figure 7a). The zonal wind amplitude in general is smaller than that of 274 the meridional winds. We also see peak height offsets between the two hemispheres. 275 The amplitudes at 95 km are displayed in Figure 7b. The increase in 2002 spring 276 The spring equinox amplitude in 2004 is larger as shown by equinox is less obvious. 277 larger area of the 25 m/s contour line. There is clear fall equinox peak during all four 278 years, but no apparent QBO related variation for the fall equinox season.

### 279 **5. Mean Zonal Winds**

#### 280 5. 1 TIDI Results

281 The mean zonal winds from a 60-day window during spring equinox are displayed in 282 Figure 8. The distinction between the two phases of QBO is very obvious. During the 283 eastward phase of QBO, there is a large region of westward zonal winds centered at the 284 equator and extended to about 100 km, whereas the westward zonal wind region is 285 smaller and wind magnitude is smaller during the westward phase of QBO. We also noticed enhancements of the eastward wind near 30°N and 40°S during the eastward 286 287 phase of the QBO. Hence, there is a further increase of the meridional gradient of the 288 mean zonal wind during the eastward phase of QBO. There is a strong eastward wind 289 region in the southern hemisphere during 2002.

### 290 5. 2 TIME-GCM Results

- 291 The simulation from the NCAR TIME-GCM for spring equinox is shown in Figure 9.
- Even though the westward wind region is much weaker and smaller, it also appears
- 293 during the eastward phase of the QBO as the TIDI observational results had shown.
- 294 The large eastward region in the southern hemisphere is present as in the case of the TIDI
- 295 observations, though with smaller wind magnitude. The TIME-GCM diurnal tide
- amplitudes range from 60 to 70 percent of that observed by TIDI.

### 297 **6. Discussions**

### 298 6.1 March Equinox Season

299 During this season, we see strong QBO dependence in TIDI data and TIME-GCM runs.

300 The westward zonal winds at the equator coincide with the enhanced the meridional and

301 zonal wind migrating tide. That should not be a surprise. The northern spring equinox 302 coincides with the peak stratospheric zonal wind magnitude (westward and eastward) of 303 the stratosphere QBO. Therefore, we should see the maximum filtering effect of the 304 gravity waves. Moreover, equinox is also the season when the migrating diurnal tide 305 peaks. Combining these two factors, we should expect the maximum QBO modulation on 306 the migrating diurnal tide during this season.

307

308 The TIME-GCM results are interesting because it reproduces the TIDI mean zonal wind 309 pattern, although to a lesser extent. At the same time the TIME-GCM diurnal tide 310 increase during the eastward phase of the QBO is also smaller (barely noticeable). It is 311 worth pointing out that both the enhancements in the tide and the westward zonal wind 312 region near the equator during the eastward phase of QBO are generated by the TIME-313 GCM based on the NCEP data at 10 hPa, while the tidal amplitudes at 10 hPa from the 314 GSWM remain the same for all four years. Compared to TIDI observations (~30 m/s), 315 the TIME-GCM mean westward winds near the equator is small (< 10 m/s), so are the 316 increase in the tide amplitude. The TIME-GCM tide amplitude is small in general. 317 That is likely because the spatial resolution (half of the scale height in vertical and 5x5 318 degrees in horizontal) of the model is not sufficient for migrating diurnal tide with 319 vertical wavelength of ~ 20 km. A new double resolution TIME-GCM 1.3 (under testing) 320 is able to produce more comparable amplitude. The TIDI mean zonal wind result can 321 be a reference for future tuning of the TIME-GCM. It will be interesting to see after 322 adjustment to the TIME-GCM model to better match the mean zonal wind, would the 323 tidal amplitude also match well.

324

### 325 6.2 Tidal Amplitudes and Mean Zonal Winds

The TIDI data show a strong connection between the tidal amplitude and mean zonal winds associated with QBO as observed in the past by HRDI. Large tidal amplitude is observed when the westward zonal wind appeared near the equator. The QBO effect on the mean zonal wind is not limited in the lower latitudes; we also noticed increase of the eastward zonal wind in the mid-latitudes during eastward phase of QBO (spring equinox).

### 333 **6.3 Inter-hemispheric Asymmetry**

334 There is a large asymmetry in tidal amplitude. Moreover, the mean zonal wind also 335 exhibits asymmetry between the two hemispheres even during spring equinox. The 336 eastward wind in the southern mid-latitudes is particularly during 2002 spring equinox. 337 The mean wind is more symmetric during the fall equinox. It is possible the QBO may 338 be asymmetric and leads to asymmetries in the mesospheric zonal winds. On the other 339 hand, the TIME-GCM results tend to be more symmetric in tidal amplitude peak altitude 340 location. The asymmetry in the model mean zonal wind has the same general features 341 observed by TIDI.

342

#### 343 **7. Summary**

344 The TIDI data provide a global view of the meridional and zonal wind migrating diurnal 345 tide and mean zonal winds. The data show strong seasonal variations of the migrating 346 tide with the strongest amplitude in spring equinox. In addition, the migrating diurnal 347 tide also is strongly affected by the QBO, the strongest effect is during the spring equinox 348 in connection with a large westward zonal wind region near the equator in the 349 mesosphere, associated with the large eastward zonal wind in the stratosphere. Given 350 that the stratospheric wind magnitudes are stronger during the spring equinox, we believe 351 that it is reasonable to expect a stronger QBO modulation during that season. The QBO 352 effect on the mean zonal wind appears to extend into the mid-latitude region by increase 353 the eastward wind during the spring equinox. The TIME-GCM driven by the NCEP data 354 at 10 hPa was able to reproduce the QBO effect on the diurnal tide and mesospheric zonal 355 wind, although to a less extend. Since the TIME-GCM is driven by tidal amplitudes 356 from the GSWM output at 10 hPa without inter-annual variations, a small mesospheric 357 QBO effect is not unexpected. We expect to see future TIME-GCM produce even 358 better agreement with observations.

359

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# 512 Figure Captions

513

### 514 Figure 1 Stratosphere Zonal Winds over Singapore

- 515 Daily stratospheric winds over Singapore from 2002 to 2005. Clear quasi-biennial
- 516 oscillation can be seen. The contour step is 5 m/s.

## 517 Figure 2 TIDI Sampling Points in Local Time

- 518 The blue track (southern) is for the coldside of TIMED satellite (away from the sun)
- 519 whereas the red track (northern) is for the warmside (towards the sun). The solid dots

520 are for daytime and hollow dots for nighttime. The ascending (nighttime) and

521 descending (daytime) nodes are plotted.

# 522 Figure 3 TIDI Tide Analysis Precision

- 523 The horizontal grid size is 6 degree and vertical is 2.5 km. Each sampling will have
- 524 about 900 raw data point from a 60 day slighting window. Precisions are estimated for
- 525 the four years during the northern spring equinox. The gradual improvement of the
- 526 precision from 2002 to 2005 is due to reduction of the ice on the TIDI optics. The
- 527 contour step is 0.5 m/s.

# 528 Figure 4 TIDI Migrating Diurnal Tide Westward one (W1) in Meridional Winds

- 529 The migrating diurnal tide amplitude altitude and latitude variations based on 60-day
- 530 period observations starting at day 51 (February 20). For easier comparison, the data
- from the same season of the four different years are plotted on the same page. The
- altitude and latitude variations during northern spring equinox are plotted in Figures 4a.
- 533 The latitude and seasonal variation at 95 km are plotted in Figure 4b. The altitude and
- seasonal variation at  $21^{\circ}$ N are shown in Figure 4c. The contour step is 10 m/s.

# 535 Figure 5 TIMEGCM Migrating Diurnal Tide W1 in Meridional Winds

- 536 Same as Figure 4 for the TIME-GCM results at the TIDI sampling points. The contour
- 537 step is 5 m/s for 4a and 4b. The color scale is different for the two plots. Figure 4c uses
- 538 contour step 10 m/s.

# 539 Figure 6 TIDI Migrating Diurnal Tide W1 in Zonal Winds

- 540 The spring equinox altitude and latitude variations of the diurnal tide amplitude are
- 541 plotted in Figure 6a (contour step 10 m/s). The latitudinal and seasonal changes at 95
- 542 km are shown in Figure 6b (contour step 15 m/s).

## 543 Figure 7 TIMEGCM Migrating Diurnal Tide W1 in Zonal Winds

- 544 Same as Figure 6 for the TIME-GCM zonal winds. The contour step for Figure 7a and
- 545 7b is 5 m/s.

## 546 Figure 8 TIDI Zonal Mean Winds

- 547 Mean zonal winds from 60-day slighting window for the northern spring equinox season
- 548 (starting on day 51, February 20). The contour step is 10 m/s.

# 549 Figure 9 TIMEGCM Zonal Mean Winds

- 550 Same as Figure 8 for the mean zonal winds from the TIME-GCM runs.
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